

Direct Search for Dark Matter with the XENON Experiment

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Exploring the Dark Sector KIAS March 16, 2015



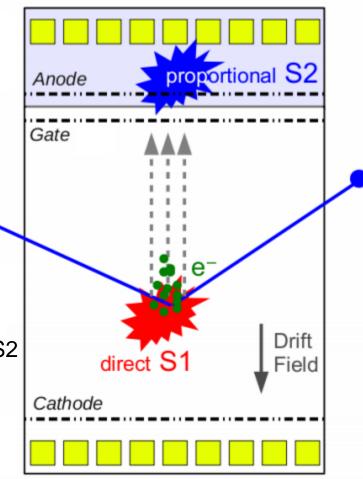
Liquid Xenon Time Projection Chambers





Particle detection with a xenon TPC

- o LXe target
- Particles hitting a Xe atom generate direct scintillation light S1 and free electrons
- Cathode and gate generate a drift field and drive the electrons upwards
- Gate and anode generate the extraction field
- Electrons reaching the liquid surface get extracted into the gas phase, accelerate and generate secondary proportional scintillation S2
- PMT arrays detect the generated signals





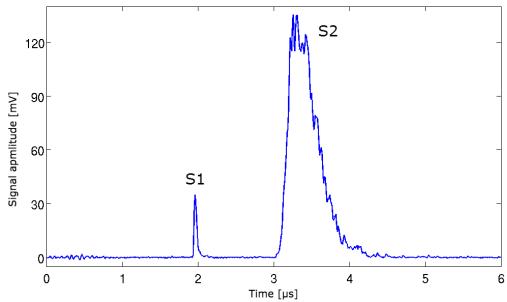
Particle interactions in liquid xenon

Two interaction types in IXe:

- Electronic recoils (γ,e)
- Nuclear recoils (n,WIMP)

$$\circ \left(\frac{S2}{S1}\right)_e > \left(\frac{S2}{S1}\right)_n$$

With a good understanding of those S2 to S1 ratios, a background reduction of >99% is possible





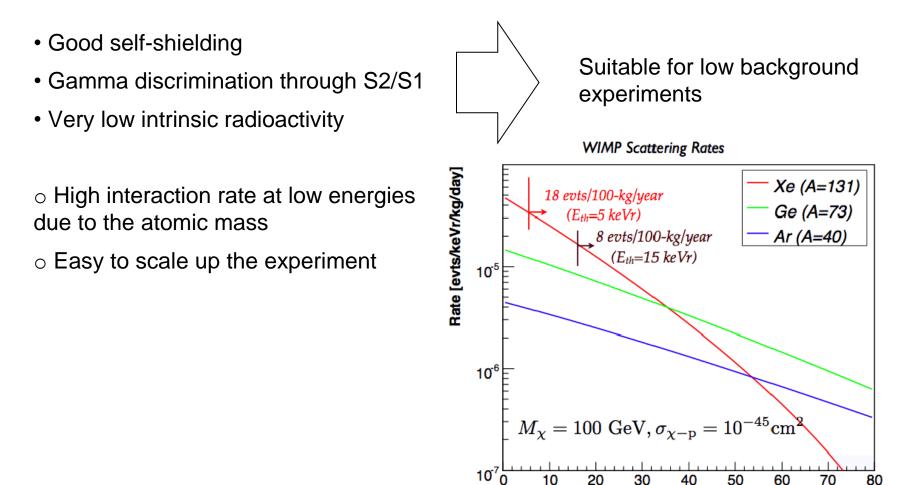
Position reconstruction in a xenon TPC

- Xenon TPC are able to achieve sub-millimeter resolution for event reconstruction
- Since the electron drift velocity at a given field is known, the vertical position can be calculated from the electron drift time
- The position in the horizontal plane is determined by comparing the signal strengths in different PMTs





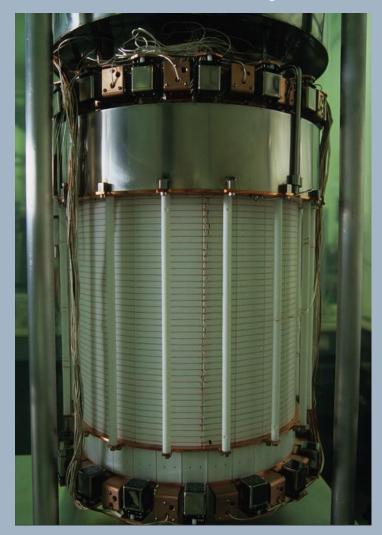
Advantages of a xenon TPC



Recoil Energy [keVr]

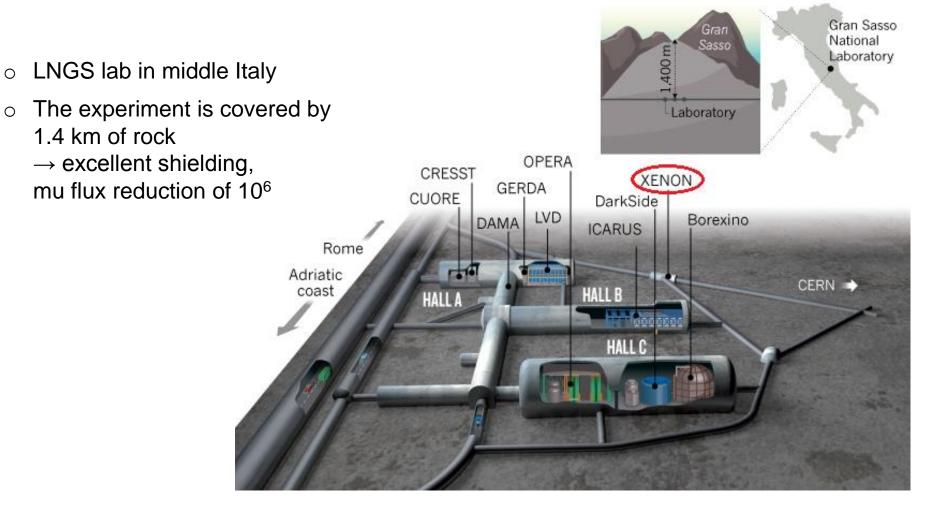


The XENON100 Experiment





Location of the XENON experiment





XENON100 specifications

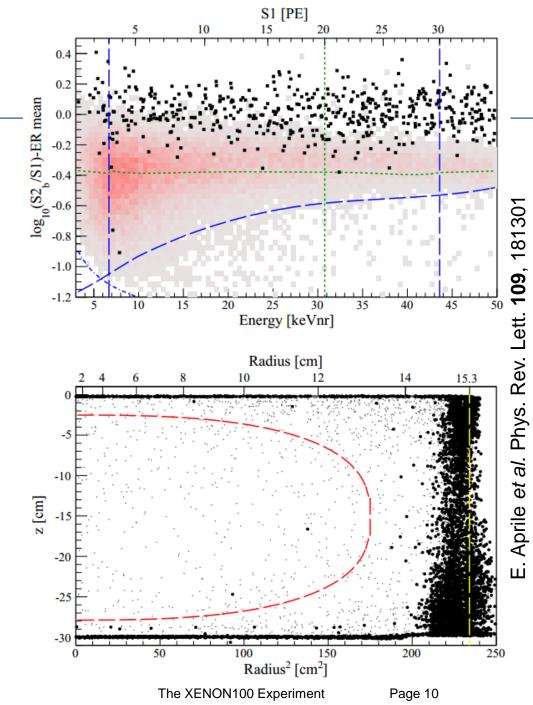
- Diameter and height of target volume ~ 30cm
- Total xenon mass ~ 160 kg
- o Target mass 62 kg
- 242 low radioactivity1" PMTs are used for signal detection
- QE of the PMTs ~ 30%
- Spatial resolution ~ 3 mm in x-y
 ~ 0.3 mm in z





XENON100 results

- 225 life days of data is evaluated
- After S2/S1 and position cuts the expected background is ~ 1 event
- 2 events are observed
 26% probability for a background fluctuation to result in the excess
- Cross section limits for SI and SD can be derived



E. Aprile et al. Phys. Rev. Lett. 109, 181301 10-39 成均館大學校 SUNGKYUNKWAN UNIVERSITY DAMA/Na XENON100 (2012) ----- observed limit (90% CL) WIMP-Nucleon Cross Section [cm²] 10-40 Expected limit of this run: CoGeNT (2012) DAMA/I $\pm 1 \sigma$ expected $\pm 2 \sigma$ expected CDMS-Si (2013) 10-41 **Limits** XENON10 (2013) 10-42 CRESST-II (2012) SuperCDMS (2014) 10^{-43} For a ~50 GeV WIMP the SI limit is 2*10⁻⁴⁵ cm^{2,} for SD neutron interaction $3.5*10^{-40}\,\text{cm}^2$ 10-44 LUX (2013) 10-45 20 40 50 60 100 200 300 400 7 8 9 1 0 1000 30 6 WIMP Mass [GeV/c²] 10⁻³⁴neutron XENON100 limit (2013) XENON100 limit (2013) proton 10⁻³⁵ ± 2σ expected sensitivity ± 2σ expected sensitivity SD WIMP-neutron cross section [cm²] 10⁻³⁵ 10⁻³⁷ 10⁻³⁸ ± 1σ expected sensitivity ± 1σ expected sensitivity 10⁻³⁶ CDM 10⁻³⁵ XENON10 PICASSO ZEPLIN-10⁻³⁷ XENON10 10⁻³⁸ IceCube (b COUPP 10 10 10¹ 10² 10^{3} 10⁴ 10¹ 10² 10³ 10⁴ WIMP Mass [GeV/c²] Page 11 WIMP Mass [GeV/c²] 3/9/2015

E. Aprile et al. Phys. Rev. Lett. 111, 181301

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Outlook Into the Future





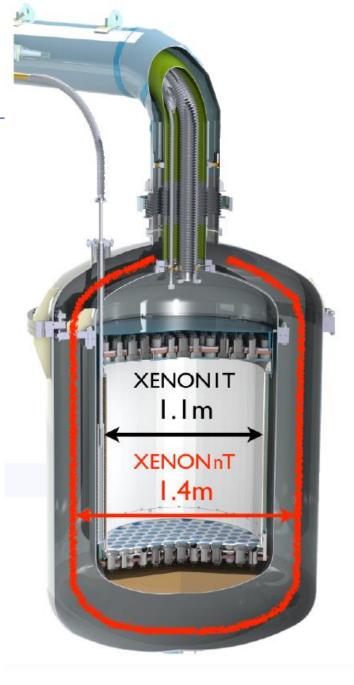
XENON1T

- Next generation XENON detector is under construction in LNGS
- Operation should start later this year
- Total xenon mass ~ 3t, target mass 1t
- Light is collected using 248 2" PMTs
- Shielded by 10 m of water and an active muon veto
- Expected sensitivity is 2 orders of magnitude better then XENON100



XENONnT

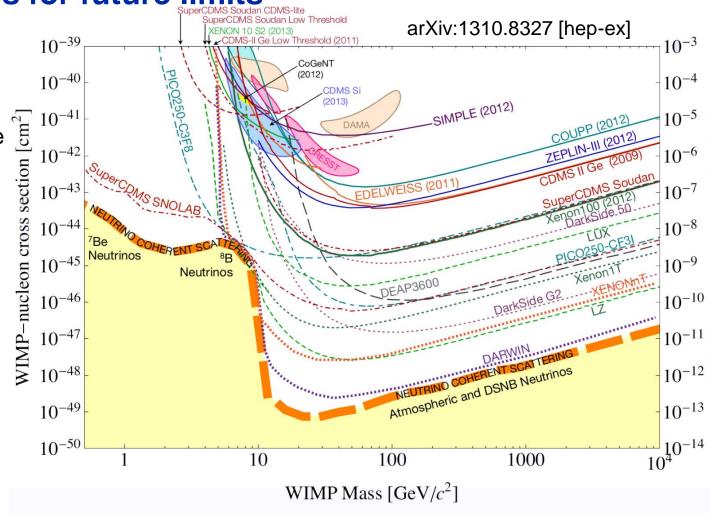
- Most XENON1T parts are designed in a way that allows for an easy upgrade
- $\circ~$ The upgrade would increase total Xe mass to ~7t $\,$
- Sensitivity would improve another order of magnitude
- o Planed for 2018





Predictions for future limits

 Experimentally accessible parameter space is limited by the neutrino background and is expected to be fully reachable within next 10-15 y





Thank you for your attention

Any questions?